

Chapter 34

The Laws of Electromagnetism Maxwell's Equations Electromagnetic Radiation Laws of Geometrical Optics

Maxwell's Equations

$ \bigoplus_{\substack{\text{closed}\\\text{surface}}} \vec{E} \bullet d\vec{A} = \frac{q_{\text{inclose}}}{\varepsilon_o} $	Gauss's Law for Electricity
$\oint_{\substack{\text{closed}\\ \text{surface}}} \vec{B} \bullet d\vec{A} = 0$	Gauss's Law for Magnetism
$\oint_{\substack{closed\\loop}} \vec{E} \bullet d\vec{s} = -\frac{d\Phi_B}{dt} = EMF$	Faraday's Law
$\oint_{\substack{\text{closed}\\\text{loop}}} \vec{B} \bullet d\vec{s} = \mu_o I_{\text{inclosed}} + \mu_o \varepsilon_o \frac{d\Phi_E}{dt}$	Ampere's Law as Modified by Maxwell

Maxwell's Equations - Static

$\oint_{\substack{\text{closed}\\\text{surface}}} \vec{E} \bullet d\vec{A} = \frac{q_{\text{inclose}}}{\varepsilon_o}$	Gauss's Law for Electricity
$ \bigoplus_{\substack{\text{closed}\\\text{surface}}} \vec{B} \bullet d\vec{A} = 0 $	Gauss's Law for Magnetism

Symmetry, but we have no magnetic monopoles.

If we had magnetic monopoles, then $\bigoplus_{\substack{closed \\ surface}} \vec{B} \bullet d\vec{A} = \mu_o \rho_{inclose}$

But, we have not detected any magnetic monopoles, so $\bigoplus_{\substack{\text{closed}\\\text{surface}}} \vec{B} \bullet d\vec{A} = 0$

Maxwell's Equations - Dynamic

$$\oint \vec{E} \cdot d\vec{s} = -\frac{d\Phi_B}{dt}$$
Faraday's Law
$$\oint \vec{B} \cdot d\vec{s} = \mu_o I_{inclosed} + \mu_o \varepsilon_o \frac{d\Phi_E}{dt}$$
Ampere's Law as Modified
by Maxwell

Again we have symmetry in the E and B field, but we lack magnetic monopoles.

If we had monopoles Faraday's Law would become

$$\oint_{loop} \vec{E} \cdot d\vec{s} = -\frac{I_{\rho}}{\varepsilon_{o}} - \frac{d\Phi_{B}}{dt} \qquad \text{Where } I_{\rho} = d\rho/dt, \text{ i.e. the flow of magnetic monopoles.}}$$

Maxwell's Equations - Coupling

Consider Maxwell's dynamic equations in vacuum with no free charges or currents.

$$\oint_{\substack{\text{closed}\\\text{loop}}} \vec{E} \bullet d\vec{s} = -\frac{d\Phi_B}{dt}, \ \Phi_B = \int \vec{B} \bullet d\vec{A}$$
(1)
$$\oint_{\substack{\text{closed}\\\text{loop}}} \vec{B} \bullet d\vec{s} = \mu_o \varepsilon_o \frac{d\Phi_E}{dt}, \ \Phi_E = \int \vec{E} \bullet d\vec{A}$$
(2)

How would a time changing **B** field effect an **E** field in vacuum (1) and how would a time changing **E** field effect an **B** field in vacuum (2)?

Maxwell's Equations - Coupling







Maxwell's Equations - Coupling



Also consider the electric field at another position, $x+\Delta x$ Assume the electric field only depends on the x-position! And that the electric field fills some space.

Maxwell's Equations - Coupling



Maxwell's Equations - Coupling











Maxwell's Equations - Coupling





Maxwell's Equations - Coupling



Maxwell's Equations - Coupling

Faraday's law predicts that as the electric field changes in space, this will induce a changing perpendicular magnetic field at that location.

$$\frac{dE}{dx} = -\frac{dB}{dt}$$

More generally,

$$\frac{\partial E_{y}}{\partial x} = -\frac{\partial B_{z}}{\partial t}$$

Where we have assumed: $\mathbf{E} = E(\mathbf{x}, t)\mathbf{j}$ and $\mathbf{B} = B(\mathbf{x}, t)\mathbf{k}$. This assumption will form *plane waves*, but in the most general case we need not make this assumption.

Maxwell's Equations - Coupling





Maxwell's Equations - Coupling

From Ampere's Law the relation between E and B is



Maxwell's Equations - Coupling

Thus from Maxwell's dynamic equations in vacuum the perpendicular E and B fields couple as

$$\frac{\partial B_z}{\partial x} = -\mu_o \varepsilon_o \frac{\partial E_y}{\partial t}$$
$$\frac{\partial E_y}{\partial x} = -\frac{\partial B_z}{\partial t}$$

Maxwell's Equations - E&M Waves

Further manipulation



Maxwell's Equations - E&M Waves

Further manipulation working towards the B field



Maxwell's Equations - E&M Waves

$\frac{\partial^2 E_y}{\partial x^2} =$	$\mu_o \varepsilon_o \frac{\partial^2 E_y}{\partial t^2}$
$\frac{\partial^2 B_z}{\partial x^2} =$	$\mu_o \varepsilon_o \frac{\partial^2 B_z}{\partial t^2}$

Wave equations!

1

Maxwell's Equations - E&M Waves

$\frac{\partial^2 E_y}{\partial x^2} = \mu_o \varepsilon_o \frac{\partial^2 E_y}{\partial t^2}$	Wave equations!
Assume, $\vec{E} = E_{\text{max}} \cos(k)$	$(x-\omega t)\hat{j}$ $k=2\pi/\lambda$ and $\omega=2\pi f$
Then, $\frac{\partial^2 E_y}{\partial x^2} = -E_{\max} \cos \theta$	$k(kx-\omega t)k^2$
and $\frac{\partial^2 E_y}{\partial t^2} = -E_{\text{max}}$	$\cos(kx-\omega t)\omega^2$
so, $-E_{\max}\cos(kx-\omega t)$	$k^2 = -\mu_o \varepsilon_o E_{\max} \cos(kx - \omega t)\omega^2$
Or $k^2 = \mu_0 \epsilon_0 \omega^2$	

Maxwell's Equations - E&M Waves



Maxwell's Equations - E&M Waves

$\frac{\partial^2 E_y}{\partial x^2} = \mu_o \varepsilon_o \frac{\partial^2 E_y}{\partial t^2}$ $\frac{\partial^2 B_z}{\partial x^2} = \mu_o \varepsilon_o \frac{\partial^2 B_z}{\partial t^2}$	The Waves (sines and/or cosines) are a solution to this Differential Equation (uniqueness theorem)!
	$\vec{E} = E_{\max} \cos(kx - \omega t)\hat{j}$ and $\vec{B} = B_{\max} \cos(kx - \omega t)\hat{k}$

Plane electromagnetic waves



Maxwell's Equations - E&M Waves

Back to $k^2 = \mu_0 \varepsilon_0 \omega^2$ Recall $k = 2\pi/\lambda$, where λ is the wavelength and $\omega = 2\pi f$, where f is the wave's frequency.

$$\left(\frac{2\pi}{\lambda}\right)^2 = \mu_o \varepsilon_o (2\pi f)^2$$

speed = $\lambda f = \sqrt{\frac{1}{\mu_o \varepsilon_o}} = c$, the speed of light

Where we have assumed the E field in the y-direction, the B field in the z-direction, and the propagation of the wave in the x-direction!

Plane electromagnetic waves



Maxwell's Equations - E&M Waves

Back to Coupling

Recall we coupled the electric and magnetic fields. We then de-coupled them to get the wave eq'n! Go back to Coupling

From Maxwell's dynamic equations in vacuum the perpendicular E and B fields couple as

$$\frac{\partial B_z}{\partial x} = -\mu_o \varepsilon_o \frac{\partial E_y}{\partial t} \qquad \vec{E} = E_{\max} \cos(kx - \omega t) \hat{j}$$
$$\frac{\partial E_y}{\partial x} = -\frac{\partial B_z}{\partial t} \qquad \vec{B} = B_{\max} \cos(kx - \omega t) \hat{k}$$

-For our goal now, these eq'ns are redundant. Use only the second one.

Maxwell's Equations - E&M Waves

$\frac{\partial E_{y}}{\partial x} = -\frac{\partial B_{z}}{\partial t}$	$\vec{E} = E_{\max} \cos(kx - \omega t)\hat{j}$
$\frac{\partial E_y}{\partial x} = -E_{\max}k\sin(kx - \omega t)$	$\vec{B} = B_{\max} \cos(kx - \omega t)\hat{k}$
$-\frac{\partial B_z}{\partial t} = -B_{\max}\omega\sin(kx-\omega t)$	

So, $-E_{\max}k\sin(kx-\omega t) = -B_{\max}\omega\sin(kx-\omega t)$

Reduces to, $E_{\max}k = B_{\max}\omega$, $\omega/k = \lambda f = c =$ speed of light

Thus,
$$E_{\text{max}} = cB_{\text{max}}$$
 Coupling!

$$\overline{E}(x,t) = \overline{j}E_{\max}\cos(kx-\alpha t)$$

$$\overline{B}(x,t) = \widehat{k}B_{\max}\cos(kx-\alpha t)$$

$$\overline{B}(x,t) = \widehat{k}B_{\max}\cos(kx-\alpha t)$$
Notes: Waves are in Phase,
but fields oriented at 90° to each other.

$$k=2\pi/\lambda.$$
Speed of wave is c= ω/k (= $f\lambda$)
At all times E=cB.

Plane electromagnetic waves

Maxwell's Equations - E&M Waves

Summary, $\frac{\partial B_{z}}{\partial x} = -\mu_{o}\varepsilon_{o}\frac{\partial E_{y}}{\partial t} \quad and \quad \frac{\partial E_{y}}{\partial x} = -\frac{\partial B_{z}}{\partial t} \quad Coupling \ eq'ns$ $\frac{\partial^{2} E_{y}}{\partial x^{2}} = \mu_{o}\varepsilon_{o}\frac{\partial^{2} E_{y}}{\partial t^{2}} \quad and \quad \frac{\partial^{2} B_{z}}{\partial x^{2}} = \mu_{o}\varepsilon_{o}\frac{\partial^{2} B_{z}}{\partial t^{2}} \quad Wave \ eq'ns$ $\vec{E} = E_{\max} \cos(kx - \omega t)\hat{j}$ $\vec{B} = B_{\max} \cos(kx - \omega t)\hat{k} \quad Wave \ sol'ns, \ propagating \ in \ the \ x - dir.$ $\frac{\omega}{k} = \lambda f = \sqrt{\frac{1}{\mu_{o}\varepsilon_{o}}} = c \quad the \ speed \ of \ light$ $E_{\max} = c \ B_{\max} \quad coupling \ of \ the \ fields$

Example

The earth is ~93 million miles from the sun. How long dose it take light leaving the sun to reach the earth?

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1 mile ~ 8/5 km 93 x 10⁶ miles x (8/5)(1000m/mile) = 1.49 x 10¹¹ m

Speed = distance/time, time = distance/speed

 $1.48 \times 10^{11} \text{ m} / (3 \times 10^8 \text{ m/s}) = 496 \text{ sec} = 8.3 \text{ min}$

Example

Express the electric and magnetic fields of a sinusoidal plane electromagnetic wave having a frequency of 3.00 GHz and traveling in the positive x-direction, if the magnitude of the maximum electric field is 300 V/m.

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$$E = E_{\max} \cos(kx - \omega t) \hat{j} \qquad E_{\max} = 300 \, \text{%}_{\text{C}}$$
$$\omega = \frac{f}{2\pi} = \frac{3 \times 10^9}{2\pi} \stackrel{\text{l}}{=} 24.77 \times 10^8 \, \frac{\text{rad}}{\text{sec}}$$
$$c = \frac{\omega}{k}, \, k = \frac{\omega}{c} = \frac{4.77 \times 10^8 \, \frac{\text{rad}}{\text{sec}}}{3 \times 10^8 \, \frac{\text{m}}{\text{sec}}} = 1.59 \,/ \text{m}$$
$$\vec{E} = 300 \, \text{\%}_{\text{C}} \cos((1.59 \,/ \text{m})x - (4.77 \times 10^8 \, \frac{\text{rad}}{\text{sec}})) \hat{j}$$

Example

Express the electric and magnetic fields of a sinusoidal plane electromagnetic wave having a frequency of 3.00 GHz and traveling in the positive x-direction, if the magnitude of the maximum electric field is 300 V/m.

$$\vec{B} = B_{\text{max}} \cos(kx - \omega t) \hat{k} \qquad B_{\text{max}} = \frac{E_{\text{max}}}{c} = \frac{300^{\circ}\%}{3 \times 10^8} = 10^{-6} \text{ T}$$
$$\vec{B} = 10^{-6} \text{ T} \cos((1.59 / \text{ m})x - (4.77 \times 10^8 \frac{\text{rad}}{\text{sec}})t) \hat{k}$$

The Electromagnetic Spectrum



Maxwell's Equations - Energy in Waves

Recall, the parallel plate capacitor stored energy in the electric field,

$$U = \frac{1}{2}CV^2$$

And the energy density, $u_E = U/Vol$.

$$u_E = \frac{1}{2}\varepsilon_o E^2$$

The energy is Stored in the E field. Similarly, E&M waves carry energy.

Maxwell's Equations - Energy in Waves

The energy density of an E&M waves do the the E field is

$$u_E = \frac{1}{2} \varepsilon_o E^2$$

From the symmetry in Maxwell's equation we can predict the energy density of the B field.

When $E \rightarrow B$ the $\varepsilon_0 \rightarrow 1/\mu_0$, so

$$u_B = \frac{1}{2} \frac{B^2}{\mu_o}$$

Maxwell's Equations - Energy in Waves

The total energy density is the sum of each.

$$u_{total} = u = u_E + u_B = \frac{1}{2}\varepsilon_o E^2 + \frac{1}{2}\frac{B^2}{\mu_o}$$

Using E=cB and $c^2=1/\epsilon_{\!_O}\,\mu_{\!_O}$

$$u = \varepsilon_o E^2 = \frac{B^2}{\mu_o}$$

Maxwell's Equations - Energy in Waves

Because light is typically of a fast frequency, the average energy density is most commonly used.

$$u_{ave} = \varepsilon_o \left\langle E^2 \right\rangle = \frac{1}{\mu_o} \left\langle B^2 \right\rangle$$

Our fields (E and B) are sinusoidal, thus the averages are: $\langle \cos^2(f(t)) \rangle = 1/2$

$$u_{ave} = \frac{\varepsilon_o E_{\max}^2}{2} = \frac{B_{\max}^2}{2\mu_o}$$

Maxwell's Equations - Energy in Waves

Thus the average energy density of a wave is given by

$$u_{ave} = \frac{\varepsilon_o E_{\max}^2}{2} = \frac{B_{\max}^2}{2\mu_o}$$

Units: Joules/m³

Maxwell's Equations - Energy in Waves

The Intensity of the wave is the amount of energy per unit area per time, i.e. Watts/m². Since the radiation is moving at the speed of light, c, the Intensity, I, is

$$I = cu_{ave}$$

Example:

Solar radiation has an intensity of about 1kW/m^2 on the earth's surface. What are the peak electric and magnetic fields in solar radiation?

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Solar radiation has an intensity of about 1kW/m^2 on the earth's surface. What are the peak electric and magnetic fields in solar radiation?

$$I = cu_{ave} = c \frac{\varepsilon_o E_{max}^2}{2} \rightarrow E_{max} = \sqrt{\frac{2I}{\varepsilon_o c}} \cong 868 \, \frac{W}{c}$$
$$E_{max} = cB_{max} \rightarrow B_{max} = \frac{E_{max}}{c} \cong \frac{868 \, \frac{W}{c}}{3 \times 10^8 \, \frac{W}{s}} \cong 2.89 \times 10^{-6} \,\mathrm{T}$$

Example:

How much electromagnetic energy is contained per cubic meter near the Earth's surface if the intensity of Sun light under clear skies is about 1000 W/m²?

$$U = u \times \text{Vol.}$$

$$I = cu_{ave}$$

$$\rightarrow u_{ave} = \frac{I}{c} \cong \frac{1000 \text{ W/m^2}}{3 \times 10^8 \text{ W/s}} \cong 3.33 \times 10^{-6} \text{ V/m^3} = 3.33 \text{ W/m^3}$$

$$U = u \times \text{Vol.} \cong 3.33 \text{ W/m^3} \times 1 \text{ m}^3 = 3.33 \mu\text{J}$$

Maxwell's Equations - Momentum and Radiation Pressure

The radiation pressure can be expressed as the Intensity by the velocity.

$$I = cu_{ave}, \quad P_{absorption} = \frac{I}{c} = u_{ave}$$

Where this is for an object which completely absorbs the radiation.

If the radiation is perfectly reflected, the pressure will double.

$$P_{reflecting} = 2u_{ave}$$

Maxwell's Equations - Momentum and Radiation Pressure

The momentum of radiation on a surface is the pressure times the surface area times the speed at which the radiation is hitting the surface.

Momentum
$$\rightarrow p_{absorption} = \frac{u_{ave}Area}{c}$$

 $p_{reflection} = \frac{2u_{ave}Area}{c}$

Example

Compare the attractive force of gravity due to the Sun on the Earth with the repulsive force of radiative pressure from Sun light.

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Compare the attractive force of gravity due to the Sun on the Earth with the repulsive force of radiative pressure from Sun light.

$$\begin{array}{ll} {\rm G}=6.67 \ {\rm x} \ 10^{-11} \ {\rm N} {\rm \cdot m^2/kg^2} & {\rm P_{absorption}} = \frac{I}{c} \\ {\rm R}_{\odot,\oplus} = 1.5 \ {\rm x} \ 10^{11} \ {\rm m} \\ {\rm M}_{\odot} = 2 \ {\rm x} \ 10^{30} \ {\rm kg} & {\rm F_{Rad}} = {\rm P_{absorption}} \ {\rm Area} = \frac{I}{c} \pi {\rm R_{\oplus}}^2 = 4.29 \times 10^8 \ {\rm N} \\ {\rm M}_{\oplus} = 6 \ {\rm x} \ 10^{24} \ {\rm kg} & {\rm I} = 1000 \ {\rm W/m^2} \\ {\rm R}_{\oplus} = 6.4 \ {\rm x} \ 10^6 \ {\rm m} & {\rm F_{Gravity}} = G \ \frac{{\rm M}_{\oplus} \ {\rm M}_{\circ}}{{\rm R}_{\circ-\oplus}^2} = 3.56 \times 10^{22} \ {\rm N} \end{array}$$



Waves from Antennas



Waves from Antennas



Waves from Antennas







Waves from Antennas



Waves from Antennas



Waves from Antennas



Waves from Antennas



Waves from Antennas



Waves from Antennas



The Inverse-Square Law:



Example:

An observer is 1.8 m from a point light source whose average power P=250 W. Calculate the rms fields in the position of the observer.

Example:

An observer is 1.8 m from a point light source whose average power P= 250 W. Calculate the rms fields in the position of the observer.

$$I = cu_{ave} = c \frac{\varepsilon_o E_{max}}{2}$$

$$I = \frac{P}{4\pi r^2}$$

$$E_{max} = \sqrt{\frac{2P}{c\varepsilon_o 4\pi r^2}} = \sqrt{\frac{P}{2c\varepsilon_o \pi r^2}}$$

$$E_{max} = \sqrt{\frac{250W}{2x3x10^8 \frac{m}{s}x8.85x10^{-12} \frac{C^2}{N \cdot m^2} \pi (1.8m)^2}}$$

$$E_{max} \equiv 68 \frac{V}{m}, \quad E_{RMS} = \frac{E_{max}}{\sqrt{2}} \cong 48.1 \frac{V}{m}, \quad B_{RMS} = \frac{E_{RMS}}{c} = 1.6x10^{-7}T$$